

Higher Spin Holography in de Sitter Space

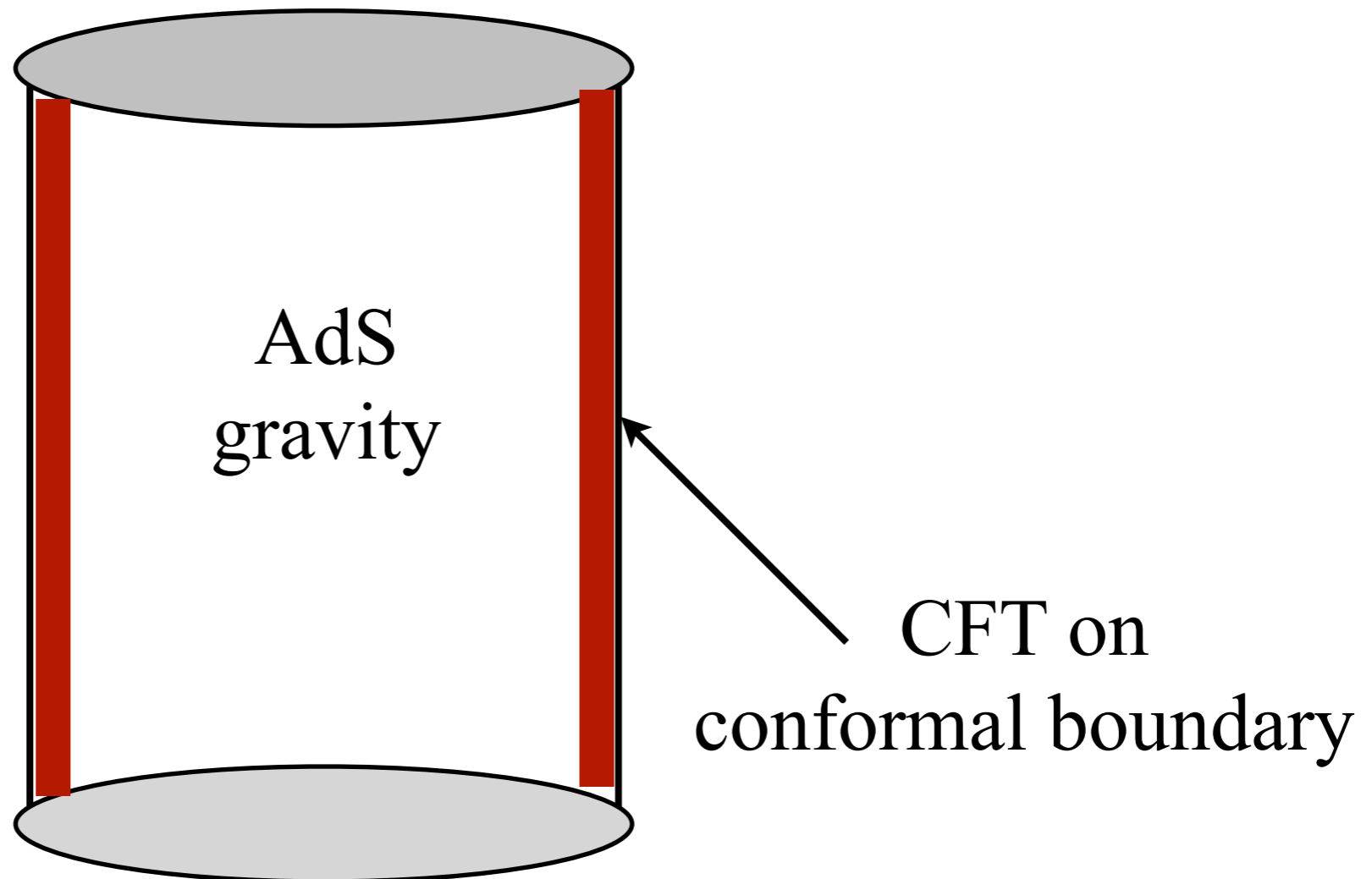
Tom Hartman
Institute for Advanced Study, Princeton NJ

Based on: 1108.5735 with D. Anninos and A. Strominger

*Ginzburg Conference
Lebedev Institute, Moscow
May 28, 2012*

Introduction

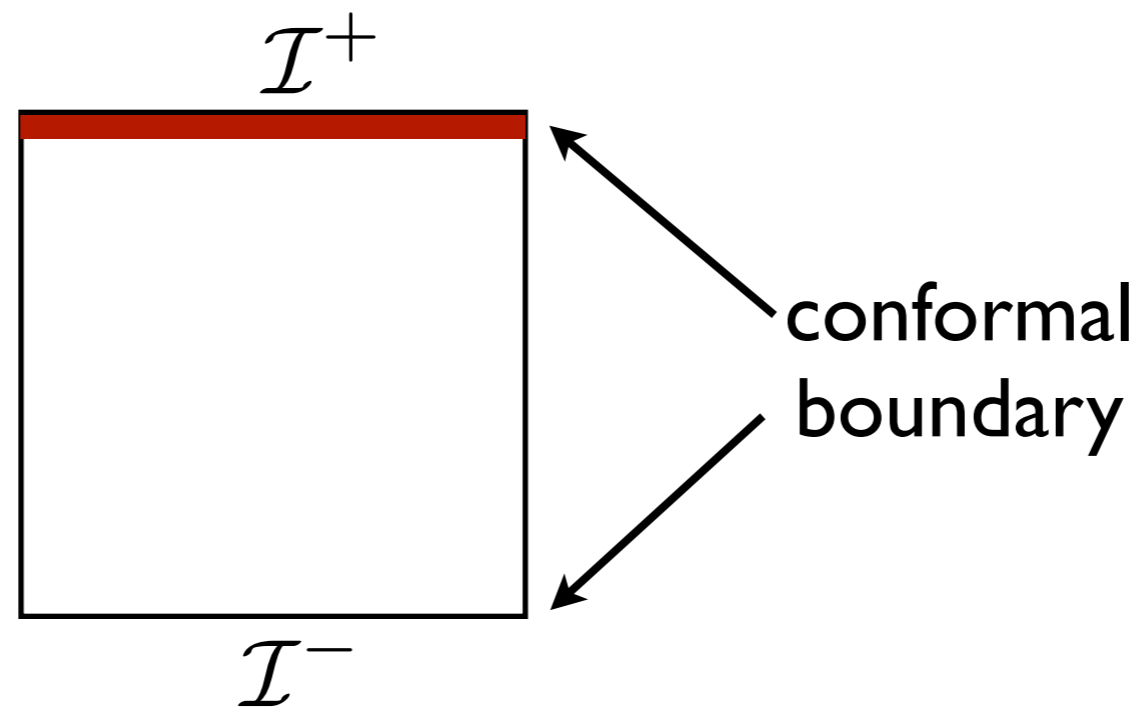
The AdS/CFT correspondence provides a non-perturbative definition of quantum gravity with negative cosmological constant.



But we live in (asymptotically) de Sitter space, in the past and future.

Introduction

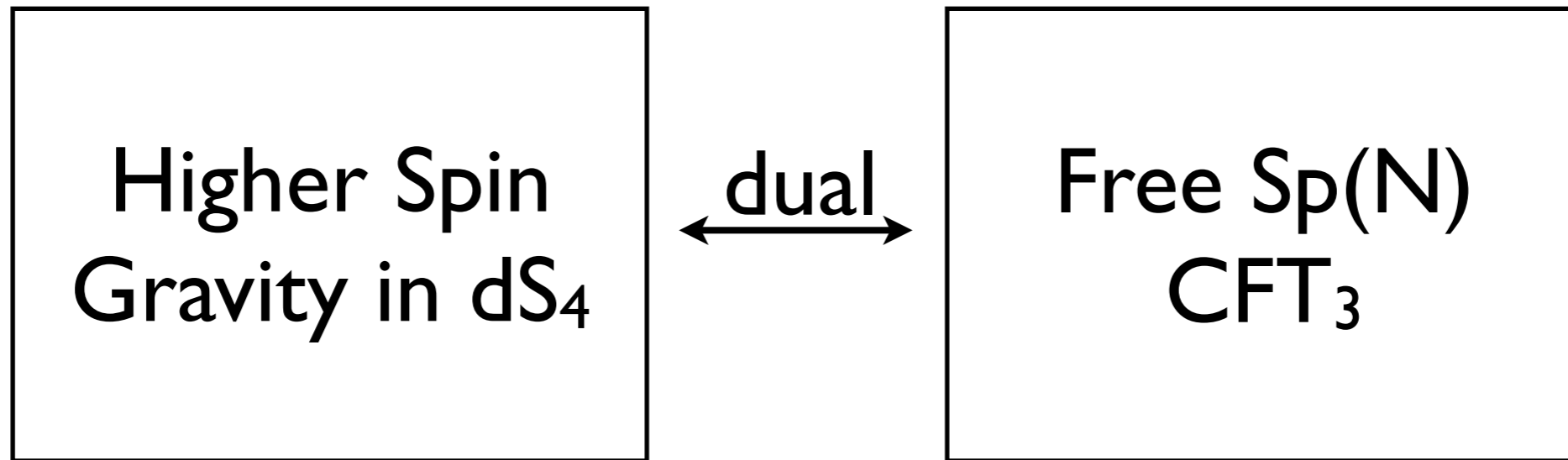
The dS/CFT correspondence, if it exists, is very different from AdS/CFT. Time is emergent:



A general dS/CFT dictionary has been proposed, but an explicit example was lacking.

Witten '01; Strominger '01; Maldacena '02.

Introduction



$$\Lambda \sim 1/N$$

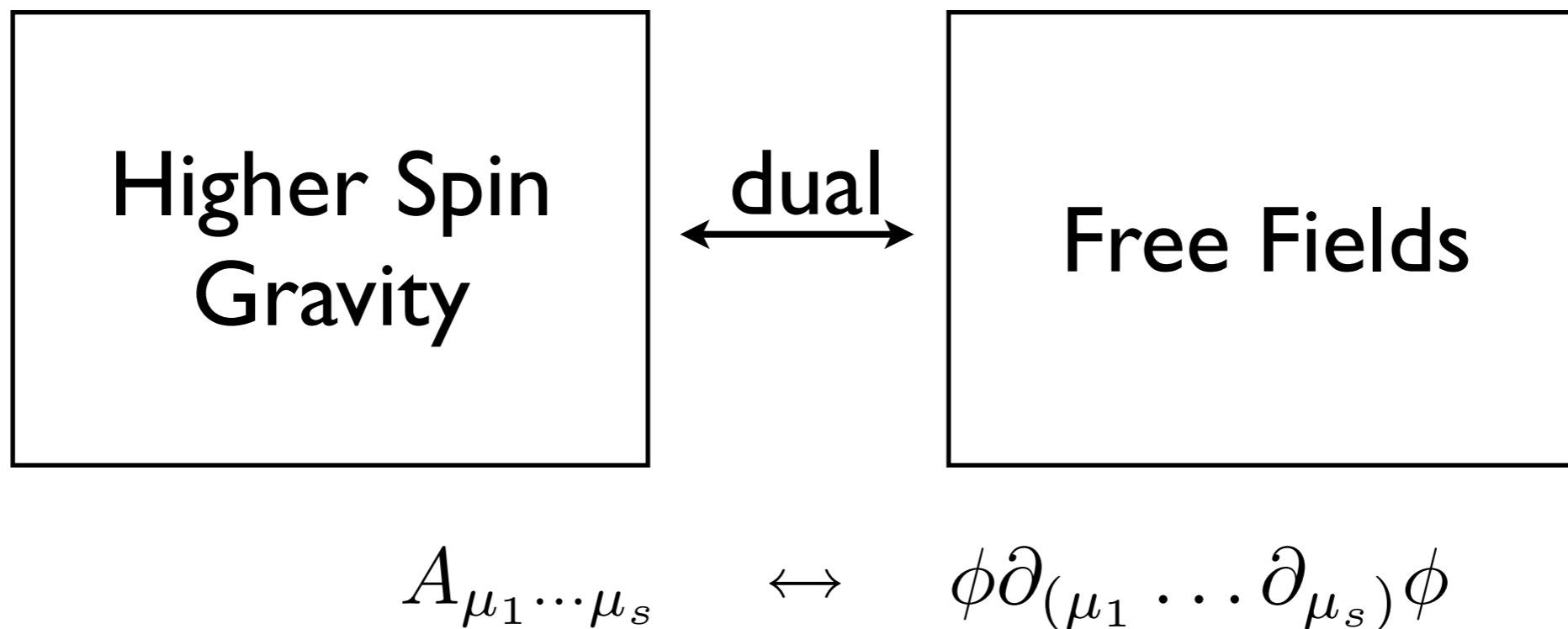
**Some Background:
Negative Cosmological Constant**

Higher Spin Dualities

4d higher spin gravity has an infinite number of massless gauge fields,

$$A_{\mu_1 \cdots \mu_s} \quad s = 0, 1, 2, 3, 4, \cdots, \quad \text{Vasiliev, 1990}$$

These gauge fields are dual to the infinite conserved currents of a free field theory.



O(N) Duality

An AdS/CFT Example

- Bulk: Vasiliev's Higher Spin Gravity
- CFT: N Free Fields -- the singlet sector of the free O(N) model

$$\phi^a \partial \dots \partial \phi^a$$

O(N) Duality

An AdS/CFT Example

- Bulk: Vasiliev's Higher Spin Gravity
- CFT: N Free Fields -- the singlet sector of the free O(N) model

$$\phi^a \partial \dots \partial \phi^a$$

Fronsdal '79

Witten

Sundborg

Mikhailov

Sezgin & Sundell

Klebanov & Polyakov

Giombi and Yin

...

O(N) Duality

An AdS/CFT Example

- Bulk: Vasiliev's Higher Spin Gravity
- CFT: N Free Fields -- the singlet sector of the free O(N) model

$$\phi^a \partial \dots \partial \phi^a$$

Fronsdal '79
Witten
Sundborg
Mikhailov
Sezgin & Sundell
Klebanov & Polyakov
Giombi and Yin
...

Evidence

- Matching spectrum
- Matching 3-point correlation functions Petkou; Giombi and Yin
- Assuming consistency: all n-point functions match
Maldacena, Zhiboedov

Proposal for $\Lambda > 0$

Note: supersymmetry was not required in AdS.

dS/CFT

The Bulk

- Vasiliev's higher spin gravity in dS_4
- Massless gauge fields of all even spins

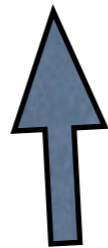
dS/CFT

The Bulk

- Vasiliev's higher spin gravity in dS_4
- Massless gauge fields of all even spins

Guessing the Boundary CFT

- Recall cosmo. constant $\Lambda \sim 1/N$
- Looking for the “O(-N)” model



this is a statement about correlators

The $\text{Sp}(N)$ Model

N free anticommuting scalars (“ghosts”)

$$S_{cft} = \frac{1}{2} \int d^3x \Omega_{ab} \partial\chi^a \cdot \partial\chi^b$$

$\Omega \equiv$ antisym. symplectic form

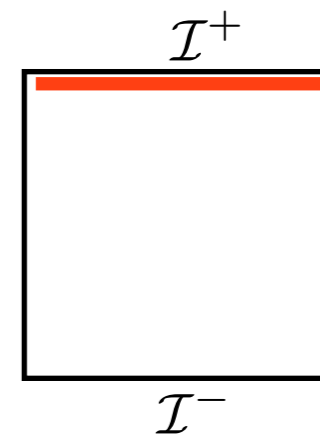
The Sp(N) Model

N free anticommuting scalars (“ghosts”)

$$S_{cft} = \frac{1}{2} \int d^3x \Omega_{ab} \partial\chi^a \cdot \partial\chi^b$$

$\Omega \equiv$ antisym. symplectic form

- In Lorentzian signature, this would be non-unitary.
- Our CFT is Euclidean; unitarity is not an issue.
- Nonetheless bulk unitarity should be encoded somehow in the CFT.



The Claim

Vasiliev gravity in $dS_4 = Sp(N)$ CFT

Evidence

- Spectrum
- 3-point correlators

Correlators

Matching Correlators

Strategy

- Relate bulk and boundary correlators to their AdS counterparts, by $N \rightarrow -N$
- Then evidence for $O(N)$ duality carries over to this case.

This is an easy way to compute, but is just for convenience. All quantities are defined intrinsically in dS and the $Sp(N)$ model, *not* by analytic continuation.

CFT Correlators

$$\langle J^{(s_1)} J^{(s_2)} \dots J^{(s_n)} \rangle_{cft} = \text{Diagram}$$

N-dependence is just an overall factor.

CFT Correlators

$$\langle J^{(s_1)} J^{(s_2)} \dots J^{(s_n)} \rangle_{cft} = \text{Diagram}$$

N-dependence is just an overall factor.

Fermi statistics:

$$\langle J \dots J \rangle_{Sp(N)} = - \langle J \dots J \rangle_{O(N)}$$

So “Sp(N) = O(-N)”

Bulk Correlators

What should we compute? (“The dictionary”)

AdS/CFT

$$Z_{cft}[X_0] = Z_{bulk}[X_0] = \int DX e^{-S_{bulk}[X]}$$

cft correlators \sim bulk vev

Bulk Correlators

What should we compute? (“The dictionary”)

AdS/CFT

$$Z_{cft}[X_0] = Z_{bulk}[X_0] = \int DX e^{-S_{bulk}[X]}$$

cft correlators \sim bulk vev

dS/CFT

$$Z_{cft}[X_0] = \Psi_{bulk}[X_0] = \int DX e^{iS_{bulk}[X]}$$

cft correlators \neq bulk vev

Bulk Correlators

What should we compute? (“The dictionary”)

AdS/CFT

$$Z_{cft}[X_0] = Z_{bulk}[X_0] = \int DX e^{-S_{bulk}[X]}$$

cft correlators \sim bulk vev

dS/CFT

$$Z_{cft}[X_0] = \Psi_{bulk}[X_0] = \int DX e^{iS_{bulk}[X]}$$

cft correlators \neq bulk vev

Bulk Correlators

How to compute it?

$$\Psi_{bulk}[X_0] \sim e^{iS_{bulk}[X]}$$

- This could be computed by solving the Vasiliev equations perturbatively in de Sitter.
- However it has already been done in AdS, **Giombi, Yin**
 - so easier to compute by analytic continuation
 - Continuing Giombi-Yin to positive CC, we find

$$\langle J \cdots J \rangle_{dS} = -\langle J \cdots J \rangle_{EAdS}$$

Correlators Match.

Therefore, borrowing AdS 3-pt functions from Giombi and Yin,

$$\langle JJJ \rangle_{dS} = \langle JJJ \rangle_{Sp(N)CFT}$$

The results of Maldacena and Zhiboedov implies that all n-point correlators match.

Final Comments

Entropy?

3d CFTs have a quantity called “F”, the partition function on a 3-sphere, which is always decreases under RG flow.

Myers et al; Jafferis; Jafferis et al; Casini & Huerta

The dS/CFT dictionary $Z_{cft} = \Psi_{bulk}$ relates this quantity to cosmological entropy:

$$\text{Entropy} = -2F$$

This relates the second law of thermodynamics to recent progress in 3d CFT.

second law \leftrightarrow F-theorem.

Problem: Compute the on-shell action of higher spin gravity on a 4-sphere.

Comments

1. Usually, analytic continuation of correlators does not give a reasonable theory; the $O(N)$ model and Vasiliev gravity are very special in this regard.
2. Non-perturbatively, $Sp(N)$ model is not “ $O(-N)$ ”
3. An interacting critical $Sp(N)$ duality also exists
in AdS: Klebanov & Polyakov
4. Time evolution = RG Flow; ex: Mass deformation

Conclusion

Higher spin gravity is dual to solvable (free in 3d) CFTs, and may allow a UV-complete holographic duality in de Sitter.

Things you might hope to compute:

- RG flows corresponding to universe production
- non-perturbative wavefunction of the universe
- de Sitter entropy
- static patch observables

Ultimately, seeking general rules for de Sitter holography.