Physics of transition scattering and some recent observations

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V. L. Ginzburg conference 2012 . V. N. Tsytovich . Physics of transition scattering .Forewords.

Most of participants have an opportunity to know V.L. personally or to work with him or at least to be impressed by the style of his wide range papers. Personally I was impressed several times- **first** as a student **in 1948** reading the first in my life scientific papers and those were the papers of V.L. (and I.E. Tamm, I.M. Frank) on Cherenkov emission (I cannot even imagine at that time that my future scientific activities will be close related with interest of V.L. and that I will have the opportunity to work with V.L.), - second in 1951 preparing my first scientific paper on synergetic effect of synchrotron and Cherenkov emission by relativistic particles (both topics have been top top activities of V.L.)-third in 1959 when I was invited to P. N. Lebedev Institute and started to give regularly presentations on famous V. L. seminar as basic for my second (after PHD) doctor degree defined in 1962 where V.L. was my general opponent, fourth – when we started to work together with V.L. in 1973, the year when V.L. invite me to teach on his chair "Problem of Physics and Astrophysics" Moscow Institute of Physics and Technology, and I nave the pleasure to work Professor in this chair up to present . The main our with V.L. theoretical work has been related to generalisation of famous work V. L.Ginzburg and I.M.Frank on Transition radiation (1946) which results in a theory of effect Transition scattering (V. L. Ginzburg and V. N.Tsytovich 1973) having at present many new applications which will be the subject of **this presentation**. This activity was summarized in 3 monographs (one of them is substantially enlarged monograph "Transition Radiation and Transition Scattering", Adam Hilger 1990)

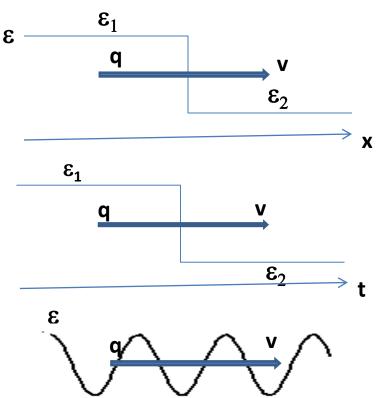
V. L. Ginzburg conference 2012 . V.N.Tsytovich General features of transition scattering physics.

Transition radiation of particle moving through boundary of two media V. L.Ginzburg and I.M.Frank (1946) .Change of self-energy. Global conservation of energy and momenta

Transition radiation of particle with sudden change of dielectric constant in time. Emission Is present for v =0

Simplest case where the dielectric constant is changed both in time and space is the wave Of dielectric constant $\varepsilon = \varepsilon_0 + \varepsilon_1 \cos(\omega t - kx)$ –transition Scattering. Emission Is present for v =0

The main problem is cross-section of transition scattering which is estimated as Q=Ze; $\sigma = Z^2 \sigma_T$ Emission exist for very heavy particles- ions . For many thermal heavy particles the Doppler width is extremely small $\Delta \omega / \omega$ about $v_T/c = (T/mc^2)^{1/2}$. For coherent structures the rate of emission is determined by square of coherent electrons



V. L. Ginzburg conference 2012 . V.N.Tsytovich General features of transition scattering physics.

The conditions for effective scattering by ions (wavelength larger than screening length) is fulfilled for rather high frequencies $\omega < \omega_{pe} c/v_T$ and specially for small angle of scattering θ and small frequency difference $\Delta \omega = |\omega_{inc} - \omega_{scat}|$

$$\frac{\Delta\omega}{\omega} < \theta^2 \cdot \frac{{}^{v}T}{c} < \frac{\omega_{pe}}{\omega} \qquad \qquad \text{or} \qquad \qquad \theta^2 \cdot \frac{{}^{v}T}{c} < \frac{\Delta\omega}{\omega} < \frac{\omega_{pe}}{\omega}$$

The effect of ion scattering in these conditions is big. Natural questions are 1) Does there exist straightforward derivation ? Yes 2) is this effect observed ? Wildly observed but sometimes not recognized

1.Field of incident wave E_{inc} excites both the current of particle oscillations j_{ocs} proportional to E_{inc} and the nonlinear current j_{nl} proportional to the field E_q of uniformly moving charge qand E_{inc} . The nonlinear responses for high frequencies are determined only by electron mass m_e and therefore the nonlinear current excites the transition scattered wave with amplitude of Thomson cross-section for scattering on free electrons. For electrons the interference of both usual and transition scattering decreases since added are not the cross-sections but amplitudes (matrix elements) and for ions the oscillating current is negligible and only transition current creates the scattered wave. The given above conditions correspond to dominations of scattering on ions and are thus confirmed by such simple calculations

V. L. Ginzburg conference 2012 . V.N.Tsytovich . Long term existing myth on difference of scattering in media on fluctuations and sum of scattering on individual particles

Theory of scattering of high frequency of electromagnetic waves on plasma fluctuations was created in 60-c E.E.Salpeter Phys. Rev., 120, 1528 (1960), 122 (1961) reviewed in excellent book M. N. Rosenbluth and N.Rostoker Phys. Fluids, 5, 776 (1962)

J.Sheffield Plasma Scattering of Electromagnetic Radiation, Academic Press, New York, San Francisco, London, 1975 and used in many experiments on plasma diagnostic by laser scattering in fusion devices and numerous other experiments on scattering of electromagnetic radiation in plasmas.

$$\begin{split} n_{0}^{(e)}S(\mathbf{k}_{-},\omega_{-}) &= 2\pi \left| 1 - \frac{G_{-}^{(e)}}{\epsilon_{-}} \right|^{2} \int \Phi_{0}^{(e)}(v^{(e)})\delta(\omega_{-} - \mathbf{k}_{-} \cdot \mathbf{v}^{(e)})d\mathbf{v}^{(e)} & \omega_{-} = \omega_{(inc)} - \omega_{(sc)}, \\ + 2\pi (Z^{(i)})^{2} \left| \frac{G_{-}^{(e)}}{\epsilon_{-}} \right|^{2} \int \Phi_{0}^{(i)}(v^{(i)})\delta(\omega_{-} - \mathbf{k}_{-} \cdot \mathbf{v}^{(i)})d\mathbf{v}^{(i)} & \mathbf{k}_{-} = \mathbf{k}_{(inc)} - \mathbf{k}_{(sc)} \\ n_{0}^{(i)} = n_{0}^{(e)}/Z^{(i)} \end{split}$$

In fact correspond to usual Thompson scattering of single electron \downarrow correspond to amplitude of transition scattering on electrons and \downarrow corresponds to transition G scattering on ions. In beginning Sheffield stated that scattered can only electrons G'and the additional terms in fluctuation

$$\epsilon(\mathbf{k},\omega) = 1 + G^{(e)}(\mathbf{k},\omega) + G^{(i)}(\mathbf{k},\omega)$$

$$f^{(e)}(\mathbf{k},\omega) = \int d\mathbf{v} \frac{4\pi e^2}{k^2(\omega - \mathbf{k} \cdot \mathbf{v} + i0)} \left(\mathbf{k} \cdot \frac{\partial \Phi^{(e)}}{\partial \mathbf{p}}\right)$$
$$f^{(i)}(\mathbf{k},\omega) = \int d\mathbf{v} \frac{4\pi (Z^{(i)})^2 e^2}{k^2(\omega - \mathbf{k} \cdot \mathbf{v} + i0)} \left(\mathbf{k} \cdot \frac{\partial \Phi^{(i)}}{\partial \mathbf{p}}\right)$$

approach are due to correlations with electrons shielding ions and \checkmark is used for that

V. L. Ginzburg conference 2012 . V.N.Tsytovich Proof that the scattering is produced also by ions

The statement that really ions are scattering can be checked by considering the influence of scattering on distributions of ions. At present the equation for ion distribution $\Phi^{(i)}$ taking into account the influence of the incident and scattered waves can be found using nonlinear currents of third order in fields including the fields created by free particles and Find the conservation laws of energy and momentum. The second term describing the transition scattering on ions satisfies the conservation laws with ion distribution which is a direct proof that the energy and momentum in this scattering process is transferred to ions.

$$\frac{dW_{sc}}{dt} + \frac{dW_{inc}}{dt} + \frac{dE^{(i)}}{dt} = 0 \quad \frac{d\mathbf{P}_{(sc)}}{dt} + \frac{d\mathbf{P}_{(inc)}}{dt} + \frac{d\mathbf{P}^{(i)}}{dt} = 0; \quad \mathbf{P}^{(i)} = \int \mathbf{p}\Phi_{\mathbf{p}}^{(i)}\frac{d\mathbf{p}}{(2\pi)^3}.$$

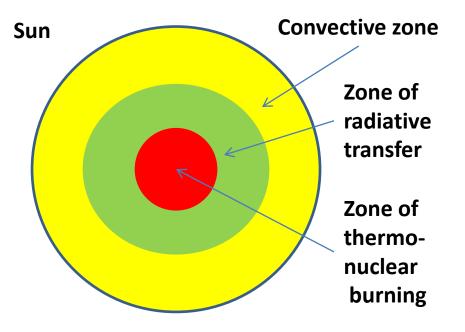
In general case with first term of scattering taking into account both electrons and ions take part in conservation laws. **The scattering in plasmas is simple a sum of scattering on electrons and ions** and the fluctuation theory only reflects this fact and with transition scattering taken into account the scattering can be described by simple physics not referring to fluctuations. This **physical approach is important** since its very rare that the particle distributions in plasmas are thermal and different effects are responsible for changes of the distribution functions.

Thus the correct interpretation deals not with correct formulations but with appropriate physics which can change the results obtained in observations

V. L. Ginzburg conference 2012 . V.N.Tsytovich Interpretation of observations

In all fusion experiments and other scattering diagnostic experiments the observers are dealing with total scattered waves including the transition scattering. The measuring of features of ion distributions is possible only because of transition scattering.

The observed resonance features of scattering in the case where \mathcal{E}_{-} is close to zero correspond to **Mandelshtamm- Brillouin** or **Raman** scattering obviously are present due to transition scattering process.



The ratio of optical to neutrino emission depends on solar opacity depending on radiative transfer which As due to radiation scattering on electrons and ions. From 20% to 60% the radiation is scattered by ions (Boerker DB). For given optical emissivity the opacity determines the temperature at the centre and therefore determined the emissivity of **high energy neutrino.**

V. L. Ginzburg conference 2012 . V.N.Tsytovich Scattering of radio waves in ionosphere and PMSE

1.Diognostic of ionosphere is usually made by scattering of radio waves similarly as diagnostic of fusion devices by lasers.

 Transition scattering can be due to massive macroscopic objects if there size is of the order or less than the wave length - emission is proportional to the square of coherent electrons in the structure. Candidates are dust particles or even dust clouds with the size less than wave length
 PMSE-Polar Mesospheric Summer Echo – Facilities EISCAT, PMVE- Winter Echo -

Main features 1)The intensity of backscattered signal from the height 80-85 km is two orders of magnitude larger

when PMSE is present

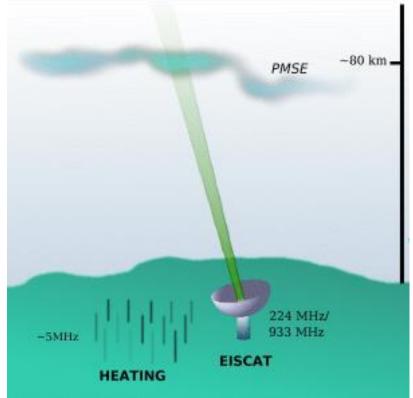
2.The frequency width of reflected signal is Very small Δω<1MHz

UHF 933 MHz

VHF 224 MHz

 $\Delta \omega / \omega < 10^{-9}$ T=(110-140)K°





<u>V.L.Ginzburg conference 2012. V.N.Tsytovich . Radar scattering from mesospheric clouds</u> <u>Strong radar backscatter from the mesopause region has</u> <u>been observed at 50 MHz</u> ¹⁹⁸¹⁻¹⁹⁸⁵ at 224 MHz ¹⁹⁸⁸ and at 933 MHz ¹⁹⁸⁹ These backscatter echoes occur in the summer and mainly at high latitudes and they have become known as "Polar Mesospheric Summer Echoes" (PMSE). Ecklund, Czechowsky, Hopps, Roetger, Havnes at al.

<u>Noctilucent clouds (NCL)</u> at this height (about 80-85 km) have been observed by centuries ago but their intensity and duration Is increases in time after industrial revolution -the effect is certainly regulated by dust in these clouds . The region is the coldest in the Earth vicinity and the grains are coated by ice. <u>PMWE</u> have been discovered only recently and can be seen only in radar scattering.

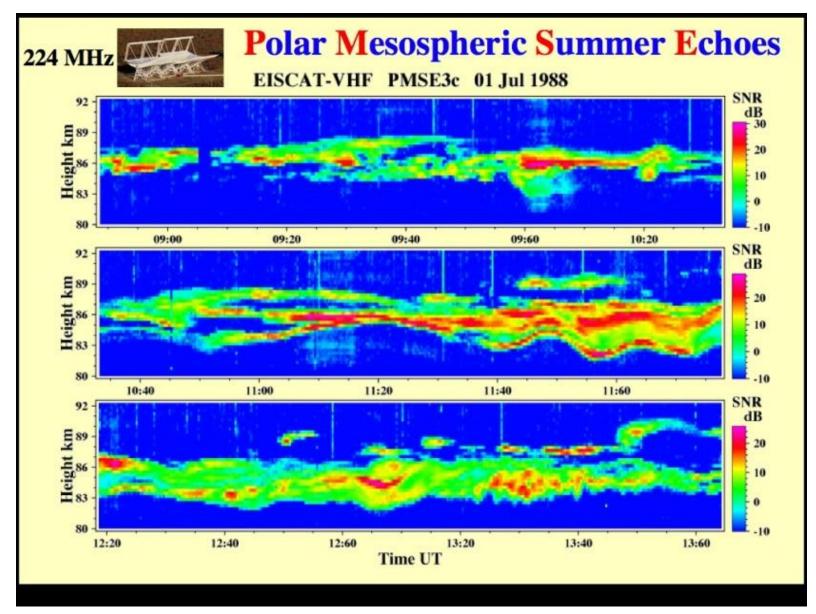
RECENT ACTIVE EXPERIMENT- Ionosphere <u>heating</u> at the PMSE and MNSE Havnes at al

Combination of high intensity of scattering and very narrow

frequency width of scattered signal <u>REQUIRES</u> mechanism of

TRANSITION SCATTERING for EXPLANATION

V.L.Ginzburg conference 2012. V.N.Tsytovich . Radar scattering MPSE



NCL- Noctilucent clouds

V.L. Ginzburg conference 2012. V.N.Tsytovich. Possible explanation of PMSE. 1.The wave length is larger than the ion Debye screening length and although in backscattering transition scattering on ions is dominant the observations cannot be explained by ion scattering a) even for low temperatures $T_i = 140K$; $\Delta \omega / \omega \sim v_{Ti} / c = 8.7 \ 10^{-7}$ **b)**enhancement of scattering is determined only by n_i and it is improbable that in presence of dust the ion density in the dust cloud region increases by factor100 **<u>2.Transition scattering by dust</u>** proposed 1990 by Havnes et all at present also <u>cannot be used for explanation.</u> The enhancement factor $Z_d^2n_d = Z_d(Z_dn_d) < Z_dn_i$ since $n_{e}=n_{i}-Z_{d}n_{d}>0$ –all electrons cannot be attached to the grains, so Z_{d} should be about **100** which implies the grain size to be larger than $(5-8)\mu$. Direct size measurements by rocked started by Havnes and by Robertson –the size are smaller. The size was

measured in active experiments by launching dust counters on rockets (Robertson 2008), but large grains are splitting by counters, the number of grains decreases with their size and cannot be properly counted, so no accurate measure of grain with largest size exist. The dust density n_d measured for 0.1μ grains is about 10^{-3} cm⁻³ The recent measurements certainly <u>exclude individual grains to</u> <u>produce the scattering enhancement by 2 orders of magnitude</u> It is believed that PMSE – are ice particles covering dust grains $0.01-0.1 \mu$; Z_d is about 3-5 and so the enhancement factor can be less 3-5

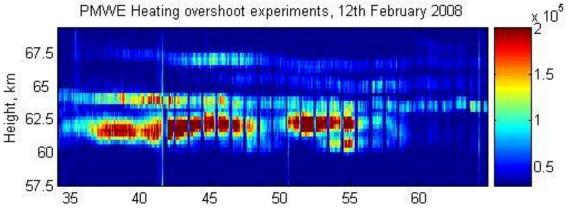
V.L. Ginzburg conference 2012. V.N.Tsytovich . Radar scattering from Polar Mesospheric Winter clouds.

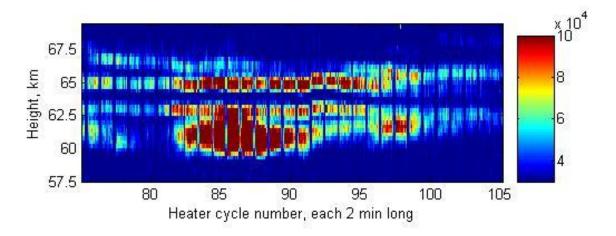
Similar arguments can be applied for interpretation of PMWE Observed 2008 PMWE Heating overshoot experiments, 12th February 2008

PMWE –probably is prodused by

very small (3-5 nm),

meteoric smoke particles (??). Can not be observed visually from ground, only with radars. The straght lines correspond to substantial decrease of scattering during heating (discussed later)Seen <u>August-May</u> – fairly rare, and weak.





V.L. Ginzburg conference 2012. V.N.Tsytovich . Most probable explanation is by transition scattering of dust clusters.

In the case when the whole dust cloud is divided by <u>dust clusters</u> the total number of **coherent electrons** Z_{cl} in the cluster can be large and transition scattering enhancement factor $Z_{cl}^2 n_{cl}$ is <u>quitesufficient</u> since similar to previous estimated for individual grains $Z_{cl} n_{cl}$ is about n_i and the enhancement factor is about Z_{cl} which for10-100 cm cluster could be much larger than 100 even for weak clustering .

Required conditions:

1.The size of the cluster should be much larger than the wave length
2.The clusters should be relative stable
<u>Present state of art</u>

1.The wave length used for radar scattering –several cm up to 3-5m can satisfy this conditions

2.Dust clumps and dust holes (dust voids) of the size about 1m have been detected in recent rocket dust detection mission to PMSE, PMWE;

3. <u>Modern general dusty plasma</u> investigations provides the <u>physics</u>, <u>laboratory</u> <u>evidence</u>, <u>theoretical</u> and <u>computational support</u> of universality of dust clump and dust void formation called as dusty plasma as <u>structurization</u>.

V.L. Ginzburg conference 2012. V.N.Tsytovich . Physics and observation of dusty plasma structurization.

Individual dust grain charges are formed by continuous by **plasma** <u>fluxes</u> on individual grains

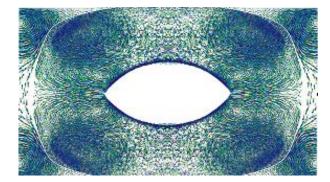
Collection of grains create collective fluxes - can confine grains, create potential well for ions and Lesage grain attraction

In linear theory homogeneous dust distribution is universally unstable

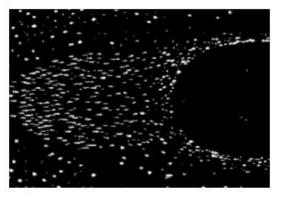
to create structures- any enhancement of dust density increases the fluxes followed by an increase of dust and ion density (Angelo1988 Morfill Tsytovich 2000).For conditions in PMSE only linear approach exist predicting maximum growth rate at (1-3)m (Tsytovich ,Havnes 2002)

Both dust **clumps** and **voids** are often observed in laboratory experiments and **on board of ISS**

Void and vortices



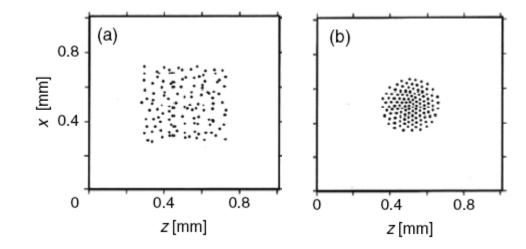
Dust clump close to dust void



V.L. Ginzburg conference 2012. V.N.Tsytovich . Boundary free clusters, numeric theoretical predictions and observations

Predicted **2001** (Khodataev et all) by PIC computations with dust Lesage attraction (due to flux shadowing)

Parameter	2D simulation	3D simulation
T_i	0.03 eV	0.03 eV
T_i	3 eV	3 eV
n	$3 \times 10^9 \mathrm{~cm^{-3}}$	$3 \times 10^9 \mathrm{~cm^{-3}}$
a	$2 \ \mu m$	$1 \mu m$
N	50	200
Z_d	5200	2600
λ_D	$23.4~\mu{ m m}$	$23.4~\mu{ m m}$
λ_M	$228~\mu{ m m}$	$228 \ \mu m$
ν_{nd}	$100 \ {\rm s}^{-1}$	$280 \ s^{-1}$
m_d	$5 \times 10^{-11} \text{ g}$	$6.3 \times 10^{-12} \text{ g}$

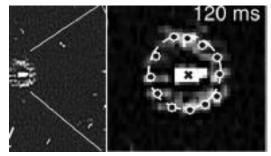


Laboratory observations 2009 (Kasachev and PK4 micro-gravity team) Formation of a Boundary-Free Dust Cluster in a Low-Pressure Gas-Discharge Plasma

Observed cluster with 12 particles cluster size 160μ

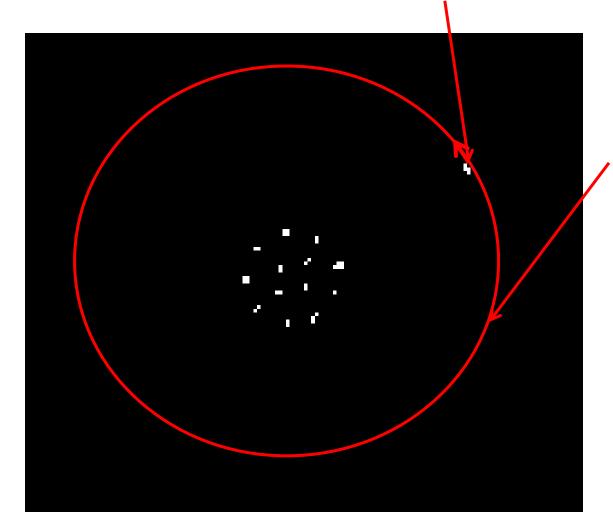
Parameters of discharges

Discharge modes	n_e, m^{-3}	T_e , eV	Z_{p1}	Z_{p2}	Grain size
dc dc-rf	$\begin{array}{c} (2\pm 0.5)\times 10^{14} \\ (4\pm 1)\times 10^{15} \end{array}$	7 ± 1 3.5 ± 0.5	19000 22000	1600 1850	a _{p1} =6µ
					Α _{p2} =0.6μ



V.L. Ginzburg conference 2012. V.N.Tsytovich . Experimental prove for existence of plasma fluxes around boundary free cluster

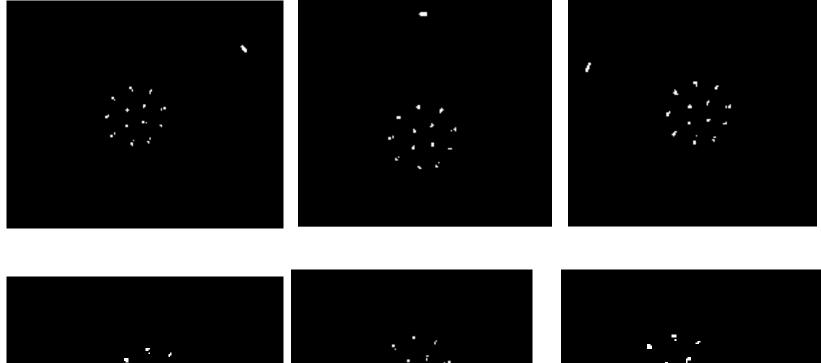
Fast particle **circulating** around boundary free cluster

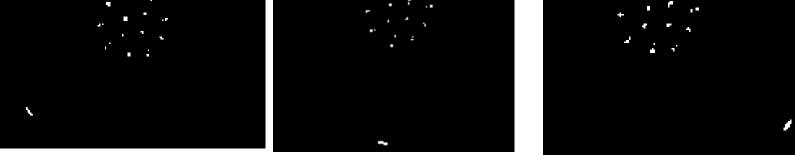


First experiments,

<u>ALLen</u> <u>Annaratone (</u>1998) Last experiments <u>Woerner</u> (2012) Trajectory of fast Grain period 3sec

V.L. Ginzburg conference 2012. V.N.Tsytovich . Experimental prove for existence of plasma fluxes around boundary free cluster

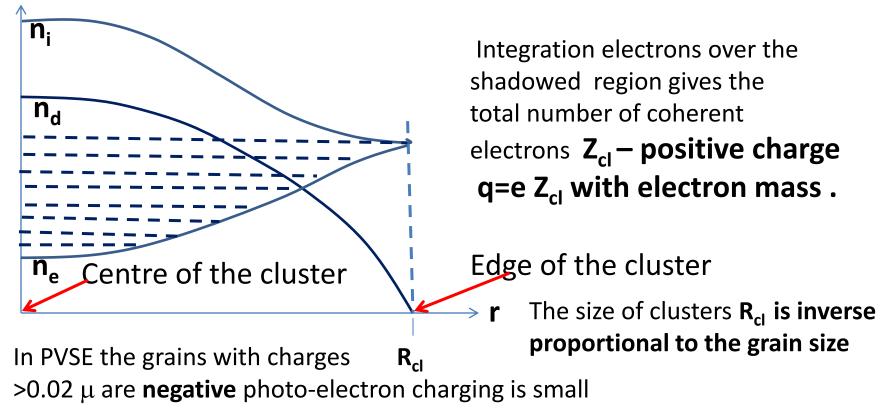




6 frames from movie of Woerner L. experiment

V.L. Ginzburg conference 2012. V.N.Tsytovich . Main features of selforganized dust clusters and their consecuencis

Typical distributions of dust ,electrons and ions obtained by numerical solutions of Master equations . Shadowed is the region of coherent <u>hole electrons</u>



Two predictions 1) for small grains the cluster size is increased and coherency can be lost (PMWE)2) sudden increase of electron temperature decreases electron density gradients ,shadowed area and decreases Z_{cl} - radar signal should decrease

V.L. Ginzburg conference 2012. V.N.Tsytovich. Active experiments on electron heating in the region of PMSE (Havnes 2003-2004)

The "overshoot" technique, an active effect found in 2003-04 to analyze the summer radar clouds(PMSE)

HF sender (Heating) with frequency 4 til 8 MHz

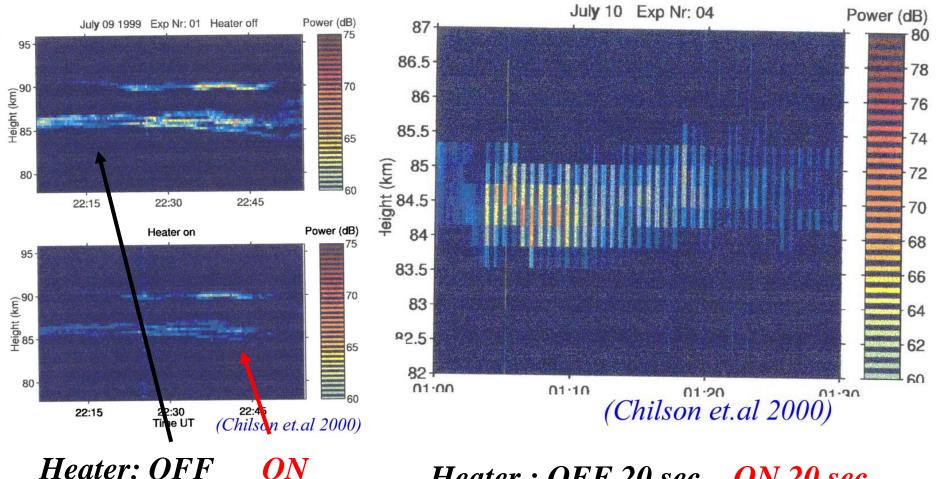


The heater affects electrons only Electrons are heated within a few milliseconds

> The electron density below the PMSE/PMWE layer control how much heater wave power which reach the PMSE layer.

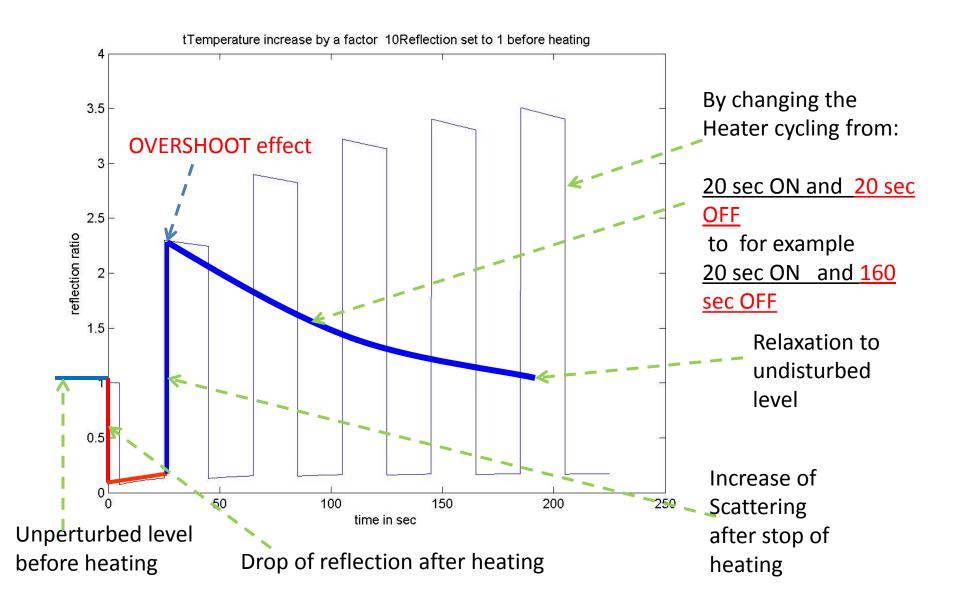
 The electron heating in the PMSE layer may be (theoretically) from negligible to around 2500 – 3000 K.

V.L.Ginzburg conference 2012. V.N.Tsytovich .Observation of Havnes O. Discovery of the electron heating effect on PMSE. Observations by the EISCAT VHF at 224 MHz



Heater : OFF 20 sec – ON 20 sec

V.L. Ginzburg conference 2012. V.N.Tsytovich. Typical active experiments results (Havnes 2004)



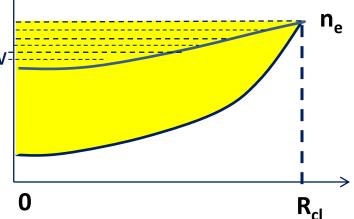
V.L. Ginzburg conference 2012. V.N.Tsytovich. Possible explanation by cluster transition scattering

1. During the heating disturbed are only electrons and other parameters of the cluster are not disturbed $T_e (dn/dr) = const$.

2.After heating (dn/dr) decreases
resulting in decrease of total number
of coherent electrons (from total yellow=
area wand to shadowed area)
3. Results – the electron heating

decreases the scattering power

4. Z_d being proportional to T_e is increases, the electron flux increases



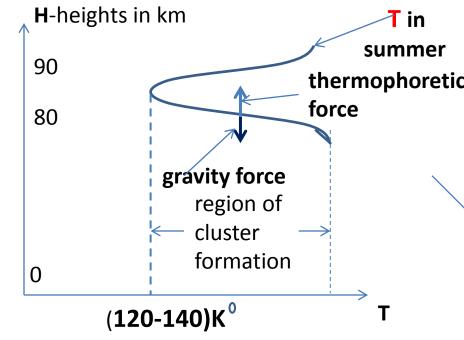
and after the heating is switched of the electrons with smaller temperature are reflected by the grains

- 5. This leads overshoot effect
- 6. The relaxation is due to ion current to the grain which requites longer time
- 7 .Estimates is this relaxation time are of the order of that observed

V.L.Ginzburg conference 2012. V.N.Tsytovich . MPSE and vorteces in mesosphere

1.Balance of gravity and thermophoretic forces- proportional to -dT/dr
2.Laboratory experiments give evidence that such balance exists but degree of ionization is much larger
3.Convection in upper atmosphere is related to system of vortices in neutral gas with velocity about several hundred meters per second- this creates a Doppler shift much less than the width of transition scattering by clusters which Is less than cm/s.

4 Planned by EC in future -A)3D rocket measutments of dust in PNSEB) Scattering facilities to be build in Norway



Transition scattering serves as a tool for Investigation of turbulence in upper atmosphere

Thank you for attention

List of references -- in Appendix

V.L.Ginzburg conference 2012. V.N.Tsytovich .

APPENDIX :References

Papers on transition radiation

1. Title Transition radiation of fast particles

authors Ginzburg V. L. and I. M. Frank I. M., Zh. Exp. Teor. Fiz., volume 16, page 15, (1946)

2. Title: THEORY OF TRANSITION RADIATION IN A NONSTATIONARY MEDIUM

Author(s): GINZBURG VL; TSYTOVIC.VN Source: ZHURNAL EKSPERIMENTALNOI I TEORETICHESKOI FIZIKI Volume: 65 Issue: 1 Pages: 132-144 Published: 1973

3. Title: FIELDS AND RADIATION OF TOROIDAL DIPOLE-MOMENTS

MOVING UNIFORMLY IN A MEDIUM

Author(s): GINZBURG VL; TSYTOVICH VN Source: ZHURNAL EKSPERIMENTALNOI I TEORETICHESKOI FIZIKI Volume: 88 Issue: 1 Pages: 84-95 Published: 1985

Papers on transition scattering general

Title: TRANSITION SCATTERING

Author(s): GINZBURG VL; TSYTOVIC.VN Source: ZHURNAL EKSPERIMENTALNOI I TEORETICHESKOI FIZIKI Volume: 65 Issue: 5 Pages: 1818-1824 Published: 1973 Times Cited: 25 (from All Databases)

 Induced resonance scattering of radiation in medium, author Tsytovich V. N. Doklady Akademii Nauk SSSR, article presented by I.E.Tamm, 54, 76 (1964);

Papers on transition scattering general (continued)

3. Title: Transition scattering in electrodynamics and general relativity Author(s): Ginzburg V.L.; Tsytovich V.N. Source: Izvestiva Vvsshikh Uchebnykh Zavedenii, Radiofizika Volume: 18 Issue: 2 Pages: 173-84 Published: 01 1975 4. Title: TRANSITION RADIATION AND TRANSITION SCATTERING IN VACUUM IN PRESENCE OF A STRONG ELECTROMAGNETIC-FIELD Author(s): GINZBURG VL; TSYTOVICH VN Source: ZHURNAL EKSPERIMENTALNOI I TEORETICHESKOI FIZIKI. Volume: 74 Issue: 5 Pages: 1621-1635 Published: 1978 5. Title: SOME PROBLEMS IN THE THEORY OF TRANSITION RADIATION AND TRANSITION SCATTERING Author(s): GINZBURG VL; TSYTOVICH VN Source: USPEKHI FIZICHESKIKH NAUK Volume: 126 Issue: 4 Pages: 553-608 Published: 1978 6. Title: ON THE DERIVATION OF THE TRANSITION RADIATION INTENSITY Author(s): GINZBURG VL; TSYTOVICH VN Source: PHYSICS LETTERS A Volume: 79 Issue: 1 Pages: 16-18 DOI: 10.1016/0375-9601(80)90303-5 Published: 1980 7. Title: Quantum theory of transition scattering Author(s): Khorev A.B.; Tsytovich V.A. Source: Soviet Physics - JETP Volume: 62 Issue: 4 Pages: 653-9 Published: Oct. 1985

V. L. Ginzburg and V. N. Tsytovich "Transition Radiation and

Transition Scattering", Nauka, Moscow, 1984, Adam Hilger, Bristol, N.Y., 1990

Transition scattering in fusion devises

 title: Plasma Scattering of Electromagnetic Radiation, Academic Press, New York, San Francisco, London, 1975 author J.Sheffield

Transition scattering in solar interior. First paper Boerker D.B. Astrop.J. v. 316 L95 (1987)

1. Title: Collective plasma effects on the solar opacity: State of the art

Author(s): Tsytovich VN; de Angelis U; Bingham R; et al. Source: JOURNAL OF QUANTITATIVE SPECTROSCOPY & RADIATIVE TRANSFER Volume: 69 Issue: 1 Pages: 9-11 DOI: 10.1016/S0022-4073(00) 00039-X Published: APR 1 2001

Title: Collective plasma processes in the solar interior and the problem

2. of the solar neutrinos deficit

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Author(s): Tsytovich V.N.; Bingham R.; de Angelis U.; et al.
Source: Physics-Uspekhi Volume: 39 Issue: 2 Pages: 103-28 Published: Feb.
1996
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3. Litle: Collective effects in quantum emission scattering

Author(s): Tsytovich VN

Source: ZHURNAL EKSPERIMENTALNOI I TEORETICHESKOI FIZIKI Volume: 109 Issue: 2 Pages: 384-392 Published: FEB 1996

4. Title: The equation of radiative transfer in the solar interior

Author(s): Tsytovich VN; Bingham R; deAngelis U; et al. Source: PHYSICA SCRIPTA Volume: 54 Issue: 3 Pages: 312-320 DOI: 10.1088/0031-8949/54/3/011 Published: SEP 1996

Transition scattering on dust ,dust clouds and dust clusters

Title: SCATTERING OF WAVES IN DUSTY PLASMAS Author(s): TSYTOVICH VN Conference: SPECIAL SESSION OF THE SPRING COLLEGE ON PLASMA PHYSICS : DUSTY PLASMAS Location: TRIESTE, ITALY Date: JUN 14-21, 1991 Sponsor(s): INT CTR THEORET PHYS; EUROPEAN ECON COMMUNITY Source: PHYSICA SCRIPTA Volume: 45 Issue: 5 Pages: 521-524 DOI: 10.1088/0031-8949/45/5/019 Published: MAY 1992 Title: ON THE ROLE OF DUST IN THE SUMMER MESOPAUSE 2 Author(s): HAVNES O; DEANGELIS U; BINGHAM R; et al. Conference: 4TH INTERNATIONAL WORKSHOP OF THE EUROPEAN INCOHERENT SCATTERING SCIENTIFIC ASSOC Location: SIGTUNA, SWEDEN Date: JUN 05-09, 1989 Sponsor(s): EUROPEAN INCOHERENT SCATTERING SCI ASSOC Source: JOURNAL OF ATMOSPHERIC AND TERRESTRIAL PHYSICS Volume: 52 Issue: 6-8 Pages: 637-643 DOI: 10.1016/0021-9169(90)90058-U Published: JUN-AUG 1990

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Thank you for attention