

Graviton Corrections to Maxwell's Equations

arXiv:1202.5800 (to appear PRD)

Katie E. Leonard and R. P.
Woodard (U. of Florida)



Classical Maxwell's Equations

$$\partial_\nu [\sqrt{-g} g^{\nu\rho} g^{\mu\sigma} F_{\rho\sigma}] = J^\mu$$

1. **Photons** ($J^\mu = 0$)

- $E^i(t, \mathbf{x}) \sim \epsilon^i e^{-ik(ct-\mathbf{x})}$

2. **Static point charge** ($J^\mu = q\delta^\mu_0\delta^3(\mathbf{x})$)

- $A_0(t, \mathbf{x}) = q/[4\pi r]$

3. **Instantaneous dipole**

- $J^0 = -\theta(t)p^i\partial_i\delta^3(\mathbf{x})$, $J^i = p^i\delta(t)\delta^3(\mathbf{x})$

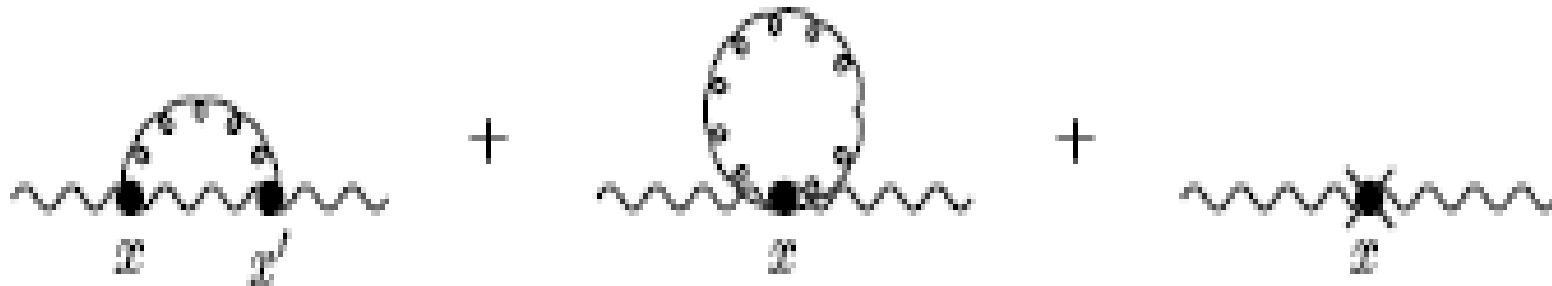
- $F_{ij}(t, \mathbf{x}) = -1/2\pi (\partial_i p_j - \partial_j p_i) \theta(t) \delta(r^2 - t^2)$



Quantum-Correcting Maxwell

- $\partial_\nu [\sqrt{-g} g^{\nu\rho} g^{\mu\sigma} F_{\rho\sigma}(x)]$
+ $\int d^4x' [\mu\Pi^\nu](x;x') A_\nu(x') = \mathcal{J}^\mu(x)$
 - 1PI 2-pt function: $i[\mu\Pi^\nu](x;x')$
 - 1PI 1-pt function: $\mathcal{J}^\mu(x)$
- Specialization to flat space
 - $[\mu\Pi^\nu](x;x') = (\eta^{\mu\nu}\partial^2 - \partial_\mu\partial^\nu) \Pi(\Delta x^2)$
 - $\partial_\nu F^{\nu\mu}(x) + \int d^4x' \Pi(\Delta x^2) \partial'_\nu F^{\nu\mu}(x') = \mathcal{J}^\mu(x)$

No Integrals at One Loop!



$$\begin{aligned}
 i[\mu\Pi\nu]_{3\text{pt}} &= \partial_{\kappa} \partial'_{\theta} \{ i\kappa V^{\mu\rho\kappa\lambda\alpha\beta} i[\Delta_{\alpha\beta}\Delta_{\gamma\delta}] i\kappa V^{\nu\sigma\phi\theta\gamma\delta} \partial_{\lambda} \partial'_{\phi} i[\Delta_{\rho\sigma}] \} \\
 &= -\kappa^2 \Gamma^2(D/2-1) / 16\pi^D (D-3)(D-2)^2 D \\
 &\quad \bullet \{ (D+1)\eta^{\mu\nu} / \Delta x^{2D} - 2D \Delta x^{\mu} \Delta x^{\nu} / \Delta x^{2D+2} \} \\
 &= \text{Same} \bullet 1/[2(D-1)] [\eta^{\mu\nu} \partial^2 - \partial^{\mu} \partial^{\nu}] 1/\Delta x^{2D-2}
 \end{aligned}$$



Schwinger-Keldysh Formalism (converting in-out to in-in)

- End of each line gets \pm polarity
- Vertices all + (**normal**) or all - (**-normal**)
- Just replace Δx^2 in propagators
 - ++ $\rightarrow ||x - x'||^2 - (|t - t'| - i\varepsilon)^2$
 - +- $\rightarrow ||x - x'||^2 - (t - t' + i\varepsilon)^2$
 - -+ $\rightarrow ||x - x'||^2 - (t - t' - i\varepsilon)^2$
 - -- $\rightarrow ||x - x'||^2 - (|t - t'| + i\varepsilon)^2$
- In-out N-pt function gives 2^N in-in N-pt



SK Effective Field Equations

- $\Pi_{\text{in-out}} \rightarrow \Pi_{++} + \Pi_{+-}$
 - $t' > t \rightarrow ++$ and $+-$ cancel
 - $\Delta r > \Delta t \rightarrow ++$ and $+-$ cancel
 - $\Delta r < \Delta t \rightarrow +- = -(++)^*$
- $\ln(\mu^2 \Delta x^2) / \Delta x^2 = 1/8 \partial^2 [\ln^2(\mu^2 \Delta x^2) - 2 \ln(\mu^2 \Delta x^2)]$
 - $\ln(\mu^2 \Delta x^2_{++}) - \ln(\mu^2 \Delta x^2_{+-}) = i\pi \theta(\Delta t - \Delta r)$
 - $\ln^2(\mu^2 \Delta x^2_{++}) - \ln^2(\mu^2 \Delta x^2_{+-})$
 $= i2\pi \theta(\Delta t - \Delta r) \ln[\mu^2(\Delta t^2 - \Delta r^2)]$
- $\Pi(x; x') = G/24\pi^2 \partial^6 \{ \theta(\Delta t - \Delta r) [\ln[\mu^2(\Delta t^2 - \Delta r^2)] - 1] \} + O(G^2)$



No Change for Photons

- $\partial_\nu F^{\nu\mu}(x) + \int d^4x' \Pi(x;x') \partial'_\nu F^{\nu\mu}(x') = \mathcal{J}^\mu(x)$
- $\partial_\nu F^{\nu\mu}(x) = 0$ works for $\mathcal{J}^\mu = 0$
- Should be true to all orders



UV Enhancement for Coulomb

- $\Phi(r) = q/4\pi r [1 + 2G/3\pi r^2 + O(G^2)]$
 - $\#G/r^2$ fixed by dimensionality
 - No dependence on finite part of counterterm
- Same sign as S-matrix techniques
 - Radkowski (1970)
 - Bjerrum-Bohr (2002)
- Opposite to RG flow from β -function
 - Robinson & Wilczek (2006)
 - Harst & Reuter (2011)
 - But other things matter



Light-Cone Pushed Out

- Point dipole p^i created at $t=0$
 - $J^0(t, \mathbf{x}) = -\theta(t) p^i \partial_i \delta^3(\mathbf{x})$
 - $J^i(t, \mathbf{x}) = p^i \delta(t) \delta^3(\mathbf{x})$
- Induced $F_{ij}(t, \mathbf{x}) = -\theta(t) (\partial_i p_j - \partial_j p_i) X$
 - 0 Loop: $\delta(r^2 - t^2)/2\pi$
 - 1 Loop: $-4G/3\pi \delta'(r^2 - t^2)/2\pi$
- Hence $\delta(r^2 - t^2 - 4G/3\pi) + O(G^2)$
 - Super-luminal but Poincare invariant



Smearing of the Light-Cone

- An old idea (Pauli 1956, Deser 1957)
 - $i\Delta[g](x;x') = 1/\{2\pi^2 \sigma[g](x;x')\} + O(\ln\sigma)$
 - $\int d^4x' \{i\Delta[g](x;x')\}^2$ doesn't exist
 - But $-\Gamma[dg] \int d^4x' \{i\Delta[g](x;x')\}^2$ might
- Something quantum gravity should do
 - Light-cone is a quantum operator
 - S-luminal propagation thru fluctuations
- We might (eventually) build starships



Summary

- $\partial_\nu F^{\nu\mu}(x) + \int d^4x' \Pi(x;x') \partial'_\nu F^{\nu\mu}(x') = J^\mu(x)$
 - $\Pi = G/(24\pi^2) \partial^6 \{ \theta(\Delta t - \Delta r) [\ln[\mu^2(\Delta t^2 - \Delta r^2)] - 1] \} + O(G^2)$
- Dynamical photons unchanged
- $\Phi(r) \rightarrow q/(4\pi r) [1 + 2G/(3\pi r^2) + O(G^2)]$
- Pulse $\rightarrow r^2 = t^2 + 4G/3\pi + O(G^2)$
 - NB Poincare invariance inevitable in pert.
- Gauge dependence requires work
 - But S-matrix results gauge independent