Dark Energy in the Local Universe

Arthur D. Chernin Sternberg Astronomical Institute Moscow University



Collaborators:

P. Teerikorpi, M.J. Valtonen, G.G. Byrd, V.P. Dolgachev, L.M. Domozhilova

GLOBAL DARK ENERGY

Einstein (1917): Cosmological constant $\Lambda \equiv$ Global antigravity, if $\Lambda > 0$

Static universe \rightarrow Gravity + Antigravity = 0

Friedmann (1922-24): Exact GR solutions with Λ

Expanding universe \rightarrow **Gravity + Antigravity = f (t)**

Riess et al. (1998), Perlmutter et al. (1999): Global acceleration

 $\Lambda > 0$



Standard ACDM Model: $DE \equiv Einstein's \Lambda$

Observations (WMAP-2006, etc.):

DE contributes 70-75% to the present total mass/energy of the universe

$$\rho_{\Lambda} = (c^2/8\pi G) \Lambda = (0.72 \pm 0.03) \cdot 10^{-29} \text{ g/cm}$$

DE antigravity dominates at

 $z < z_{\Lambda} \approx 0.7$; $t > t_{\Lambda} \approx 7$ Gyr

VACUUM

MACROSCOPIC DESCRIPTION: Einstein-Gliner vacuum

Gliner (1965):

- * Equation of state $\mathbf{p}_{\Lambda} = \rho_{\Lambda}$
- * Perfectly uniform
- * Density ρ_{Λ} is constant in any reference frame

WMAP (2006): $p_{\Lambda}/\rho_{\Lambda} = -1 \pm 0.1$

Planck (launched 14 apr 2009): ±0.03-0.02

ANTIGRAVITY

GR: effective gravitating density

.

$$\rho_{eff} = \rho + 3 p$$
DE: $\rho_{eff} = -2 \rho_{\Lambda} < 0 \rightarrow antigravity$

PHYSICAL NATURE???

MICROSCOPIC STRUCTURE:

Severe challenge to fundamental physics

Zeldovich (1967): $\Lambda \equiv \text{Quantum Vacuum}$ Arkani-Hamed et al. (2001): $\rho_{\Lambda} \sim (M_{FW}/M_{P})^{8} \rho_{P}$

 $M_{EW} \sim 1 \text{ TeV}, M_P \sim 10^{15} \text{ TeV}$

 $\rho_p \sim M_p^4$

LOCAL DARK ENERGY

DE is discovered at near-horizon distances of ~ 1 000 Mpc

Does DE act on relatively small local scales?

In principle, yes -- if $DE = \Lambda$

Specifically:

* How strong may local DE antigravity be?

* May DE antigravity dominate locally?

LOCAL GRAVITY-ANTIGRAVITY

Schwarzschild-de Sitter static spacetime:

point-like mass on DE background

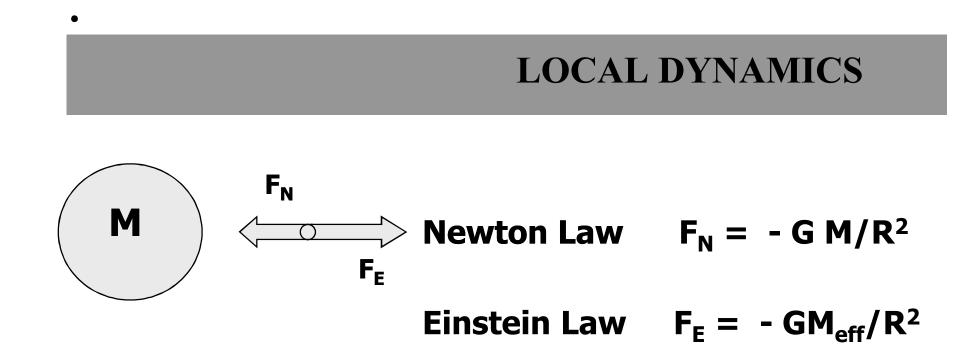
 $ds^2 = A(R) dt^2 - R^2 d \Omega^2 - A^{-1} dR^2$

A (R) = $1 - 2GM/R - (8\pi G/3) \rho_{\Lambda} R^2$

Newtonian limit:

 $1 + U \approx A^{1/2} \approx 1 - GM/R - (4\pi G/3) \rho_{\Lambda} R^{2}$

 $F(R) = - \text{grad } U = - GM/R^2 + (8\pi G/3) \rho_{\Lambda} R$

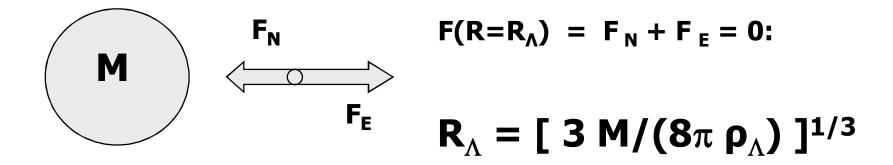


$$M_{eff} = (4\pi/3) \rho_{eff} R^3 = (4\pi/3) (\rho + 3p) R^3 = -(8\pi/3) G \rho_{\Lambda} R^3$$

 F_E = + (8π/3) G $ρ_A$ R

(per unit mass)

ZERO-GRAVITY RADIUS

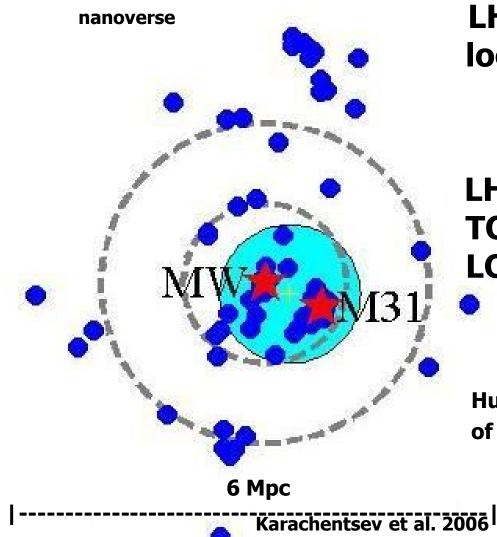


 $\approx 1 [M/10^{12} M_{sun}]^{1/3} Mpc$ (Chernin et al. 2000)

Groups of galaxies:M = (1-10) $10^{12} M_{sun} \rightarrow R_{\Lambda} = 1-2$ MpcClusters of galaxies:M = (1-10) $10^{14} M_{sun} \rightarrow R_{\Lambda} = 5-10$ Mpc

 R_{Λ} is local counterpart of global redshift $z_{\Lambda} \approx 0.7$

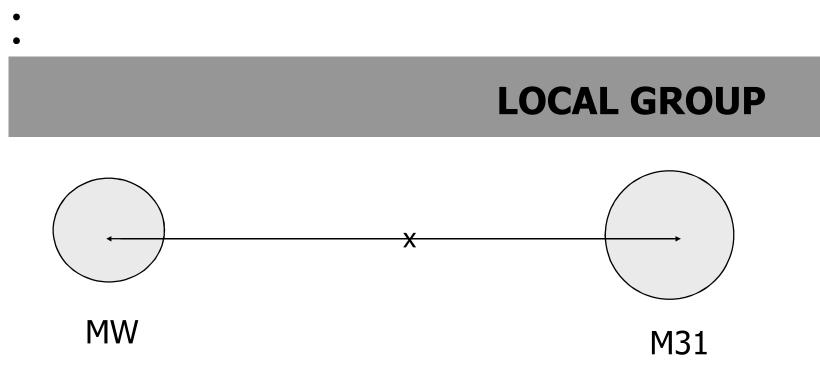
LOCAL HUBBLE CELL (LHC)



LHC = Local Group + local expansion outflow

LHC IS NATURAL SETUP TO DETECT AND MEASURE LOCAL DARK ENERGY

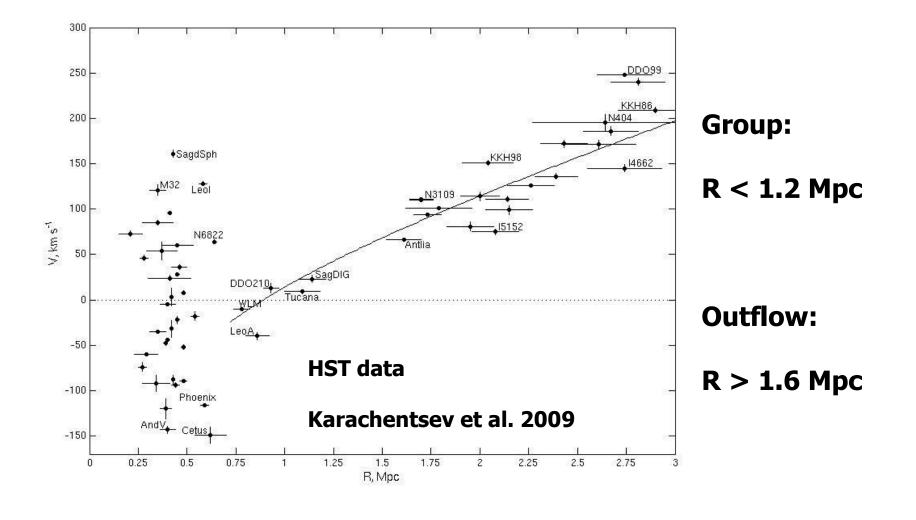
Hubble cells are typical population of the Local Universe (Hubble 1936)



Giant MW-M31 binary and \approx 50 dwarf galaxies

MW and M31 are separated by distance of 0.7 Mpc and move toward each other with relative velocity – 120 km/s now

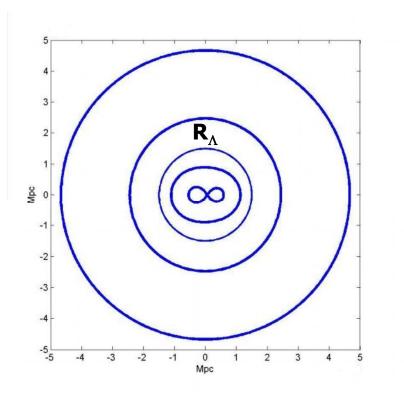
LHC PHASE PORTAIT



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LHC ZERO-GRAVITY RADIUS

LHC gravity-antigravity potential



Gravitationally bound group

 $R < R_{\Lambda}$: gravity dominates

Expansion outflow

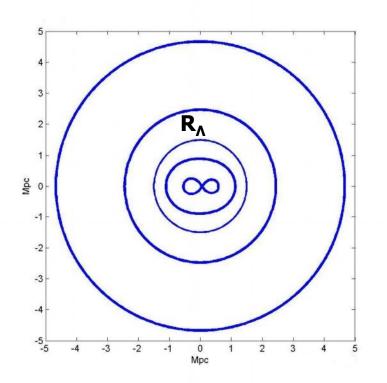
 $R > R_{\Lambda}$: antigravity dominates

Antigravity makes outflow cool: linear velocity-distance relation with low dispersion

(Chernin et al. 2000-04, Maccio' et al. 2005, Sandage et al. 2006)

LHC MODEL

LHC gravity-antigravity potential



Group:

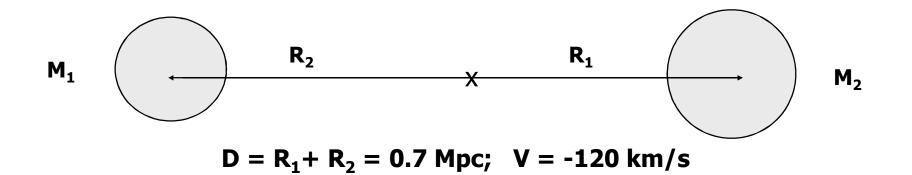
MW_M31 binary as bound linear two-body system

Expansion outflow:

dwarf galaxies as test particles moving in a spherical gravity-antigravity static potential

The cell is embedded in the uniform DE background





Linear two-body problem on DE background: equations of motion

$$d^{2}R_{1}/dt^{2} = -G M_{2}/D^{2} + G (8\pi/3)\rho_{\Lambda} R_{1}$$

 $- d^2 R_2/dt^2 = G M_1/D^2 - G (8\pi/3)\rho_{\Lambda} R_2$

(barycenter reference frame)

GRAVITY-ANTIGRAVITY POTENTIAL

 $d^2D/dt^2 = -~GM/D^2 ~+~G(8\pi/3)\rho_\Lambda~D$ $(M = M_1 + M_2~,~D = R_1 + R_2)$ The first integral

(1/2) V² = GM/D + G (4 π /3) ρ_{Λ} D² + E, (E =Const)

Effective potential : $U = -GM/D - G(4\pi/3) \rho_{\Lambda} D^2$

BINDING CONDITION

Binary is gravitationally bound, if

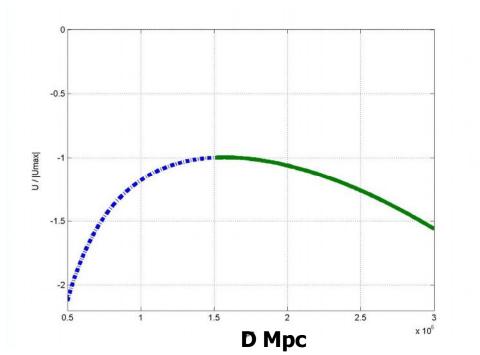
 $E < U_{max} = -(3/2) GM^{2/3} [(8\pi/3) \rho_{\Lambda}]^{1/3}$

Binding condition + first integral

 $(1/2) V^2 = GM/D + G (4\pi/3) \rho_{\Lambda} D^2 + E$

lead to lower limit to mass

 $M > M_1 = 3.3 \ 10^{12} M_{sun}$



 $(KW: M_1 = 1.1)$

TIMING ARGUMENT

Gravitational instability in Λ CDM model is terminated when antigravity becomes stronger than gravity at t = t_A = 7 Gyr (e.g. Chernin et al. 2003)

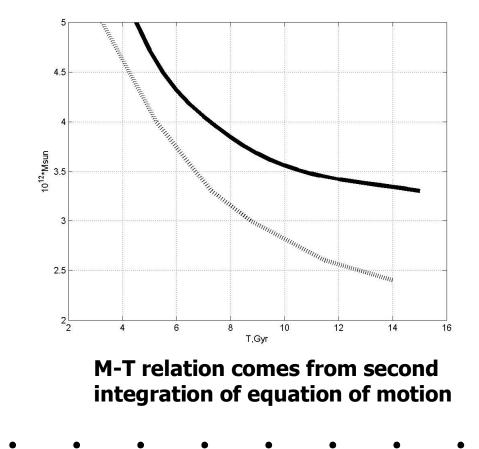
Therefore binary collapse time

$$T > t_0 - t_\Lambda \approx 7 Gyr$$

M-T relation leads to upper limit to mass:

$$M < M_2 = 4.1 \ 10^{12} M_{sun}$$

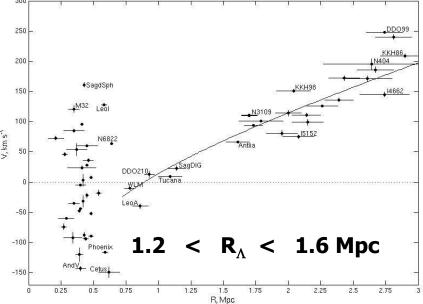
(KW: M₂ = 3.2)



LIMITS TO ZERO-GRAVITY RADIUS

 $\mathbf{R}_{\Lambda} = \left[\begin{array}{cc} 3 \text{ M} / (8 \pi \rho_{\Lambda}) \right]^{1/3} \rightarrow \mathbf{M} = (8 \pi / 3) \rho_{\Lambda} \mathbf{R}_{\Lambda}^{3}$

 $M_{3} < M < M_{4}$ $M_{3} = 1.7 \ 10^{12} M_{sun}$ $M_{4} = 3.9 \ 10^{12} M_{sun}$ $M_{3} = 10^{12} M_{sun}$ $M_{4} = 3.9 \ 10^{12} M_{sun}$



HST data: Karachentsev et al. 2009

LOCAL GROUP MASS

Four limits to group mass in combination

$3.3 < M < 3.9 \ 10^{12} M_{sun}$

If local DE density = global value ρ_{Λ}

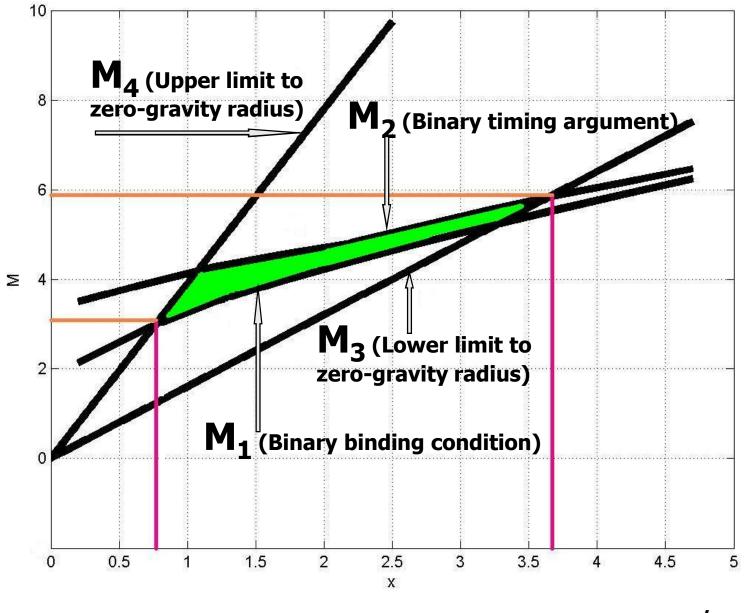
LOCAL DE DENSITY

If local DE density is unknown ρ_x , four mass limits are functions of local DE density, or $x = \rho_x / \rho_A$

- $M > M_1(x), M < M_2(x),$
- $M < M_3(x), \qquad M_4 > M_4(x)$

Observational data: D = 0.7 Mpc, V = -120 km/s

1.2 < R_{Λ} < **1.6** Mpc



 $\mathbf{x} = \rho_{\mathbf{x}} / \rho_{\Lambda}$

CONCLUSIONS

LHC is natural setup for measuring Local Group mass and local DE density at ~ 1 Mpc in self-consistent way

Total (dark matter + baryons) mass of Local Group

 $\begin{array}{rcl} 3.1 \ < \ M \ < \ 5.9 \ \ 10^{12} \ \ M_{sun} \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & \\ & & & & \\ & & &$

 Λ CDM Millennium Simulations (2008): 1.7 < M < 5.1 10¹² M_{sun}

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CONCLUSIONS

DARK ENERGY EXISTS ON SCALE ~ 1 Mpc

ANTIGRAVITY IS STRONG ON THIS SCALE

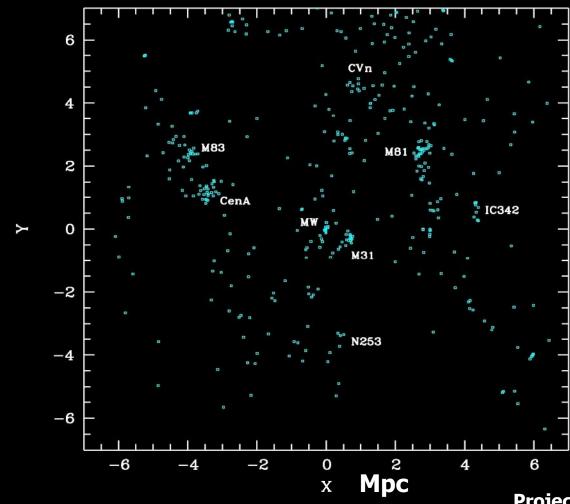
Local density of dark energy

 $0.8 < \rho_x < 3.7 \rho_\Lambda$

Local DE density at R \sim 1 Mpc is close, if not exactly equal, to global DE density at R \sim 1 000 Mpc

NEW INDEPENDENT EVIDENCE FOR EINSTEIN'S IDEA OF UNIVERSAL ANTIGRAVITY

LOCAL UNIVERSE



Karachentsev et al. (2006):

HST map of the local volume of 14 Mpc across

Projection on the Supergalactic plane

LOCAL UNIVERSE

The Local Universe is a network of receding partly overlapping "Hubble cells"

The Hubble cell is a gravitationally bound galaxy system (group or cluster) + an expansion flow of galaxies around it

The network (100-300 Mpc across) is embedded in the uniform dark energy background

KAHN-WOLTJER MODEL

$$d^2D/dt^2 = -G M/D^2 \rightarrow \frac{1}{2} V^2 = GM/D + E$$

 $(dD/dt = V; D = R_1 + R_2; M = M_1 + M_2)$

The first integral and the bound condition E < 0 imply:

$$M > M_{min} = \frac{1}{2} V^2 D/G = 10^{12} M_{sun}$$

 $M_{min} \sim 10 M_L$

First evidence for dark matter in MW and M31